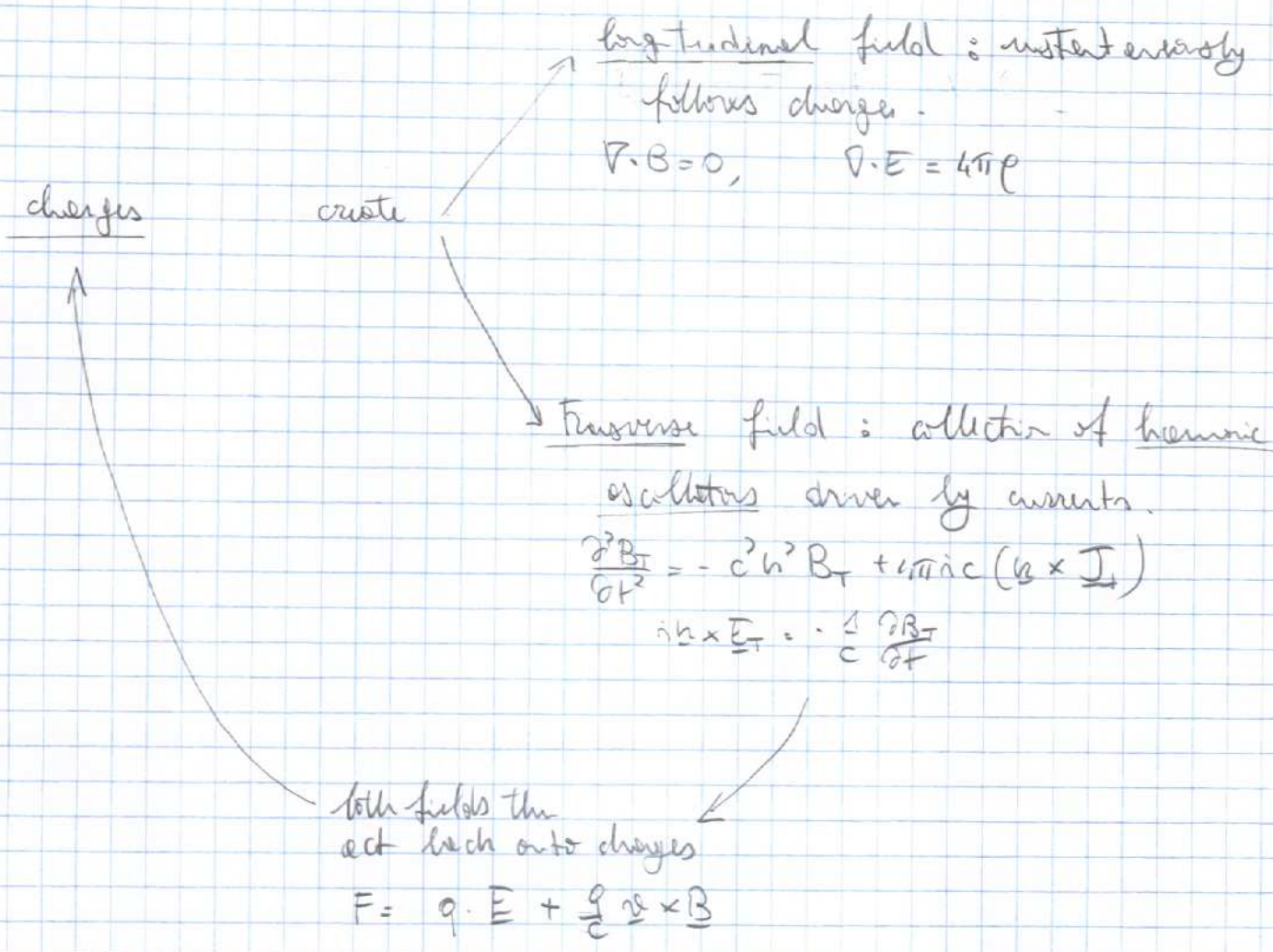


Lecture 7 : quantization of the e.m. field



Degrees of freedom :

- charge motion
- transverse e.m. field :  
 collection of harmonic oscillators  
 1 h.o. per  $\mathbf{k}$  vector in reciprocal space.

Quantization

- naive approach : replace h.o. variables with operators
- Lagrangian / Hamiltonian approach

Reference book:

C. Cohen-Tannoudji, J. Dupont-Roc, & Grynberg  
 "Introduction to QED",

Standard Lagrangian of E.D.:

$$\begin{aligned} \mathcal{L} &= \sum_{\alpha} \frac{1}{2} m \dot{r}_{\alpha}^2 + \frac{1}{8\pi} \int d^3r [\underline{E}(r)^2 - \underline{B}(r)^2] \\ &+ \sum_{\alpha} \left[ \frac{q_{\alpha}}{c} \dot{r}_{\alpha} \cdot \underline{A}(r_{\alpha}) - q_{\alpha} V(r_{\alpha}) \right] = \\ &= \mathcal{L}(r_{\alpha}, \dot{r}_{\alpha}; \underline{A}(r), \dot{\underline{A}}(r), V(r), \dot{V}(r)) \end{aligned}$$

in  $\mathcal{L}$ :  $E(r) = -\nabla V(r) - \frac{1}{c} \dot{\underline{A}}(r)$

$$B(r) = \nabla \times \underline{A}(r)$$

interaction term can be rewritten as:

$$\int d^3r \left[ \frac{1}{c} \underline{j}(r) \cdot \underline{A}(r) - \rho(r) \cdot V(r) \right]$$

Covariant form of  $\mathcal{L}$ :

$$\mathcal{L} = -\frac{c^2}{16\pi} F_{\mu\nu} F^{\mu\nu} - \sum_{\alpha} q_{\alpha} dx_{\alpha}^{\mu} A_{\mu} - \sum_{\alpha} m_{\alpha} c^2 dt_{\alpha}$$

in Fourier space :

$$d^4 = \frac{1}{2} \omega_\alpha \dot{x}_\alpha^2 + \int \frac{d^3n}{(2\pi)^3} \left\{ \frac{1}{4\pi} \left[ |\tilde{E}(n)|^2 - |\tilde{B}(n)|^2 \right] + \right.$$

$$\left. + \frac{1}{c} \left( \tilde{j}^*(n) \cdot \tilde{A}(n) + \tilde{A}^*(n) \cdot \tilde{j}(n) \right) - \tilde{\rho}^*(n) \tilde{U}(n) + \right.$$

$$\left. - \tilde{U}^*(n) \tilde{\rho}(n) \right\}$$

NOTE :

- $\tilde{A}(n), \tilde{U}(n)$  are complex variables.
- reality of  $A(z), U(z) \implies \tilde{A}(-n) = \tilde{A}^*(n); U(-n) = U^*(n)$
- fields in half reciprocal space are independent  
 $\hookrightarrow$  integral  $f$  in  $\mathcal{L}$  limited to half-space.
- real and imaginary parts of  $A(n), U(n)$  are independent real variables

A parenthesis : complex coordinates in lagrangian

$\mathcal{L}(x_1, x_2; \dot{x}_1, \dot{x}_2)$  generic lagrangian (real)

$X =$  change of variables  $(x_1, x_2) \rightarrow (X, X^*)$

$$X = \frac{1}{\sqrt{2}}(x_1 + i x_2), \quad X^* = \frac{1}{\sqrt{2}}(x_1 - i x_2)$$

generic  $f(x_1, x_2) : df = \frac{\partial f}{\partial x_1} dx_1 + \frac{\partial f}{\partial x_2} dx_2$

in complex coordinates:

17,4

$$df = \frac{\partial f}{\partial X} dX + \frac{\partial f}{\partial X^*} dX^*$$

$$\text{with } \begin{cases} \frac{\partial}{\partial X} = \frac{1}{\sqrt{2}} \left( \frac{\partial}{\partial x_1} - i \frac{\partial}{\partial x_2} \right) \\ \frac{\partial}{\partial X^*} = \frac{1}{\sqrt{2}} \left( \frac{\partial}{\partial x_1} + i \frac{\partial}{\partial x_2} \right) \end{cases}$$

$$\text{indeed } \frac{\partial}{\partial X} X = \frac{1}{\sqrt{2}} \left( \frac{\partial}{\partial x_1} - i \frac{\partial}{\partial x_2} \right) \frac{(x_1 + i x_2)}{\sqrt{2}} = 1$$

$$\frac{\partial}{\partial X} X^* = \frac{1}{\sqrt{2}} \left( \frac{\partial}{\partial x_1} - i \frac{\partial}{\partial x_2} \right) \frac{(x_1 - i x_2)}{\sqrt{2}} = 0.$$

Conjugate momentum:

$$P = \frac{\partial \mathcal{L}}{\partial \dot{X}^*} = \frac{1}{\sqrt{2}} \left( \frac{\partial \mathcal{L}}{\partial \dot{x}_1} + i \frac{\partial \mathcal{L}}{\partial \dot{x}_2} \right) = \frac{1}{\sqrt{2}} (p_1 + i p_2)$$

Lagrange equations

$$\frac{d}{dt} \frac{\partial \mathcal{L}}{\partial \dot{X}^*} - \frac{\partial \mathcal{L}}{\partial X^*} = 0 \quad \Rightarrow \quad \frac{d}{dt} \frac{1}{\sqrt{2}} (p_1 + i p_2) - \frac{1}{\sqrt{2}} \left( \frac{\partial \mathcal{L}}{\partial x_1} + i \frac{\partial \mathcal{L}}{\partial x_2} \right) = 0$$

$$\text{separating real and imaginary parts: } \frac{d}{dt} \frac{\partial \mathcal{L}}{\partial \dot{x}_{1,2}} - \frac{\partial \mathcal{L}}{\partial x_{1,2}} = 0$$

Hamiltonian

$$H = \dot{x}_1 p_1 + \dot{x}_2 p_2 - \mathcal{L} = \dot{X} P^* + \dot{X}^* P - \mathcal{L}$$

Commutators

$$[\tilde{X}, \tilde{P}] = \left[ \frac{1}{\sqrt{2}} (x_1 + i x_2), \frac{1}{\sqrt{2}} (p_1 + i p_2) \right] = 0$$

$$[\tilde{X}, \tilde{P}^\dagger] = i \neq 0$$

Back to E.D.

$$\begin{aligned} \mathcal{L} = & \sum_{\alpha} \frac{1}{2} m_{\alpha} \dot{z}_{\alpha}^2 + \int \frac{d^3 n}{(2\pi)^3} \cdot \left\{ \frac{1}{4\pi} \left[ \left| \frac{\dot{\tilde{A}}(n)}{c} + i n \tilde{U}(n) \right|^2 + \right. \right. \\ & \left. \left. - |n \times \tilde{A}(n)|^2 \right] + \frac{1}{c} (j^{\dagger}(n) A(n) + j(n) \cdot A^{\dagger}(n)) + \right. \\ & \left. - \rho(n) U^{\dagger}(n) - \rho^{\dagger}(n) U(n) \right\} = \end{aligned}$$

$$= \sum_{\alpha} \frac{1}{2} m_{\alpha} \dot{z}_{\alpha}^2 + \frac{q_{\alpha}}{c} \cdot \dot{z}_{\alpha} \cdot A(z_{\alpha}) - q_{\alpha} U(z_{\alpha}) +$$

$$+ \int \frac{d^3 n}{(2\pi)^3} \cdot \frac{1}{4\pi} \left[ \left| \frac{\dot{\tilde{A}}(n)}{c} + i n \tilde{U}(n) \right|^2 - |n \times A(n)|^2 \right]$$

eq. motion for particles

$$\frac{d}{dt} \frac{\partial \mathcal{L}}{\partial \dot{z}_{\alpha}} - \frac{\partial \mathcal{L}}{\partial z_{\alpha}} = 0$$

$$\frac{d}{dt} \left( m \dot{z}_{\alpha} + \frac{q_{\alpha}}{c} A(z_{\alpha}) \right)_i - \frac{q_{\alpha}}{c} \dot{z}_{\alpha, j} \cdot \nabla_i A_j + q_{\alpha} \nabla_i U = 0$$

$$m \ddot{z}_{\alpha, i} = -q_{\alpha} \nabla_i U - \frac{q_{\alpha}}{c} \left[ \frac{\partial A}{\partial t} + \dot{z}_{\alpha, j} \nabla_j A_i - \dot{z}_{\alpha, j} (\nabla_j A_i) \right]$$

$\underbrace{\hspace{10em}}_{\frac{d}{dt} A(z_{\alpha})}$

$$m \ddot{x}_i = q \alpha \left[ -\nabla_i U - \frac{1}{c} \frac{\partial A_i}{\partial t} \right] + \frac{q \alpha}{c} \dot{x}_j \times (\nabla \times A)_i$$

↳ Lorentz force

in fact:

$$\begin{aligned} \dot{x}_\alpha \times (\nabla \times A)_i &= \epsilon_{ijk} \dot{x}_{\alpha j} \epsilon_{nlm} \frac{\partial}{\partial x_l} A_m = \\ &= \epsilon_{nij} \epsilon_{nlm} \dot{x}_{\alpha j} \frac{\partial}{\partial x_l} A_m = \\ &= \dot{x}_{\alpha j} \left( \frac{\partial}{\partial x_j} A_j \right) - \dot{x}_{\alpha j} \frac{\partial}{\partial x_j} A_i \end{aligned}$$

eq. motion for field: i) scalar potential  $U$

$$\frac{\partial \mathcal{L}}{\partial U^*} = -\frac{1}{4\pi} \dot{\underline{h}} \cdot \left( \frac{\dot{\underline{A}}}{c} + \dot{\underline{h}} U \right) = e$$

$$\frac{\partial \mathcal{L}}{\partial \dot{U}^*} = 0$$

$$\frac{d}{dt} \frac{\partial \mathcal{L}}{\partial \dot{U}^*} - \frac{\partial \mathcal{L}}{\partial U} = 0 \quad \text{fixes instantaneous value of } U$$

$$4\pi e = \dot{\underline{h}} \cdot \left( -\frac{1}{c} \dot{\underline{A}} - \dot{\underline{h}} U \right)$$

$$\text{i.e. } \dot{\underline{h}} \cdot \underline{E} = 4\pi e \quad (\text{without eq.})$$

ii) vector potential  $A$

$$\frac{\partial \mathcal{L}}{\partial \underline{A}^*} = \frac{j(\underline{h})}{c} + \frac{1}{4\pi} \underline{k} \times (\underline{h} \times \underline{A})$$

$$\frac{\partial \mathcal{L}}{\partial \dot{\underline{A}}^*} = \frac{1}{4\pi c^2} (\dot{\underline{A}} + \dot{\underline{h}} \cdot \underline{U})$$

$$\frac{d}{dt} \left( \frac{1}{4\pi c} \left( \dot{\mathbf{A}} + ikU \right) \right) = \frac{j(n)}{c} + \frac{1}{4\pi} n \times (n \times \mathbf{A})$$

$$\frac{1}{c} \frac{d}{dt} \left( -\frac{1}{c} \dot{\mathbf{A}} - ikU \right) = -\frac{4\pi}{c} j(n) + ik \times (\underbrace{ik \times \mathbf{A}}_B)$$

$E$

$$ik \times B = \frac{1}{c} \frac{\partial E}{\partial t} + \frac{4\pi}{c} j \quad \rightarrow \text{Maxwell eq.}$$

The two other Maxwell's eqs:

$$\nabla \cdot B = 0 \quad \text{automatic from } B = \nabla \times A$$

$$\nabla \times E = -\frac{1}{c} \frac{\partial B}{\partial t} \quad \text{automatic from } E = -\nabla U - \frac{1}{c} \frac{\partial A}{\partial t}$$

$$\begin{aligned} \text{in fact: } \nabla \times \left( -\nabla U - \frac{1}{c} \frac{\partial A}{\partial t} \right) &= -\cancel{\nabla \times (\nabla U)} - \frac{1}{c} \nabla \times \frac{\partial A}{\partial t} = \\ &= -\frac{1}{c} \frac{\partial}{\partial t} \nabla \times A = -\frac{1}{c} \frac{\partial B}{\partial t} \end{aligned}$$

Simplify the Lagrangian

\*  $U$  is not a real degree of freedom, as it is related to  $A$  and particle position. Replacing in  $\mathcal{L}$  its expression  $U = \frac{1}{n^2} \left( ik \frac{\dot{\mathbf{A}}}{c} + 4\pi e \right)$  gives:

$$\begin{aligned} \mathcal{L} &= \sum_{\alpha} \frac{1}{2} m_{\alpha} \dot{\mathbf{r}}_{\alpha}^2 - \int \frac{d^3n}{(2\pi)^3} \frac{4\pi}{n^2} |e(n)|^2 + \frac{1}{4\pi} \int \frac{d^3n}{(2\pi)^3} \left[ \frac{1}{c^2} |\dot{\mathbf{A}}_{\perp}|^2 - (k \times \mathbf{A}_{\perp}^{\dagger}) (k \times \mathbf{A}_{\perp}) \right] \\ &\quad + \int \frac{d^3n}{(2\pi)^3} (j^{\dagger} \mathbf{A}_{\perp} + \mathbf{A}_{\perp}^{\dagger} j) + \frac{d}{dt} \int \frac{d^3n}{(2\pi)^3} \frac{ik}{n} [e \mathbf{A}_{\parallel}^{\dagger} - e^{\dagger} \mathbf{A}_{\parallel}] \end{aligned}$$

\*  $A_{\parallel}$  only appears in total derivative term that can be eliminated

$$\Rightarrow \mathcal{L}(\mathbf{r}_\alpha, \dot{\mathbf{r}}_\alpha; \mathbf{A}_\perp, \dot{\mathbf{A}}_\perp; \dot{\mathbf{A}}_\parallel, \dot{\mathbf{A}}_\parallel)$$

\* Coulomb term can be rewritten as

$$-\frac{1}{2} \sum_{\alpha, \alpha'} \frac{q_\alpha q_{\alpha'}}{|\mathbf{r}_\alpha - \mathbf{r}_{\alpha'}|} = - \int \frac{d^3r}{(2\pi)^3} \frac{4\pi}{n^2} |\rho(\mathbf{r})|^2 + \text{self energy of particles}$$

For real space:

$$\begin{aligned} \mathcal{L} = & \sum_{\alpha} \frac{1}{2} m \dot{\mathbf{r}}_\alpha^2 - \frac{1}{2} \sum_{\alpha, \alpha'} \frac{q_\alpha q_{\alpha'}}{|\mathbf{r}_\alpha - \mathbf{r}_{\alpha'}|} + \frac{1}{8\pi} \int d^3r \left[ \frac{1}{c^2} \dot{\mathbf{A}}_\perp^2 - (\nabla \times \mathbf{A}_\perp)^2 \right] + \\ & + \sum_{\alpha} \frac{q_\alpha}{c} \dot{\mathbf{r}}_\alpha \cdot \mathbf{A}_\perp(\mathbf{r}_\alpha) = \mathcal{L}(\mathbf{r}_\alpha, \dot{\mathbf{r}}_\alpha, \mathbf{A}_\perp, \dot{\mathbf{A}}_\perp) \end{aligned}$$

+ longitudinal field  $\rightarrow$  Coulomb energy

\* transverse field  $\rightarrow$  radiative effects

NOTE: even if the "COULOMB GAUGE"  $\nabla \cdot \mathbf{A} = 0$  ( $A_\parallel = 0$ ) is not a convenient concept, the physics of the e.m. field is described exactly. If needed, a relativistic description of particles can be included



## Conjugate momenta

$$\Pi(\mathbf{n}) = \frac{\partial \mathcal{L}}{\partial \dot{A}(\mathbf{n})} = \frac{1}{4\pi c^2} \dot{A}(\mathbf{n})$$

$$p_\alpha = \frac{\partial \mathcal{L}}{\partial \dot{x}_\alpha} = m \dot{x}_\alpha + \frac{q_\alpha}{c} A(x_\alpha)$$

$$H = \int \frac{d^3n}{(2\pi)^3} \left[ \Pi^\dagger(\mathbf{n}) \dot{A}(\mathbf{n}) + \Pi(\mathbf{n}) \dot{A}^\dagger(\mathbf{n}) + p_\alpha \dot{x}_\alpha \right] - \mathcal{L} =$$

$$= \sum_\alpha \int \frac{1}{2m_\alpha} \left( p_\alpha - \frac{q_\alpha}{c} A(x_\alpha) \right)^2 + \frac{1}{2} \sum_{\alpha\alpha'} \frac{p_\alpha p_{\alpha'}}{|x_\alpha - x_{\alpha'}|} +$$

$$+ \frac{1}{4\pi} \int \frac{d^3n}{(2\pi)^3} \left[ c^2 |4\pi \Pi(\mathbf{n})|^2 + \kappa^2 |A(\mathbf{n})|^2 \right]$$

\* light-matter interaction via minimal coupling  
replacement  $P \rightarrow P - \frac{q}{c} A$

→ field evolution → quadratic  $H$  in  $X, P \Rightarrow$  collection of h.o.

\* for every  $k$  in half-space

→ 2x polarization states

→ real + imaginary part of  $A, \Pi$

Upon quantization:

$$[A(\mathbf{n}), \Pi(\mathbf{n}')] = 0, \quad [A(\mathbf{n}), \Pi(\mathbf{n}')^\dagger] = i \kappa (2\pi)^3 \delta^3(\mathbf{n} - \mathbf{n}')$$

Define

$$\text{for } k \in \text{H.S.} \quad \begin{cases} \alpha(k) = \sqrt{\frac{1}{8\pi\hbar c k}} \left[ k A(k) + \frac{4\pi i c}{k} \Pi(k) \right] \\ \alpha^\dagger(k) = \sqrt{\frac{1}{8\pi\hbar c k}} \left[ k A^\dagger(k) - \frac{4\pi i c}{k} \Pi^\dagger(k) \right] \end{cases}$$

to obtain a full basis, has to introduce also:

$$\beta(k) = \sqrt{\frac{1}{8\pi\hbar c k}} \left[ k A^\dagger(k) + \frac{4\pi i c}{k} \Pi^\dagger(k) \right]$$

$$\text{and } \beta^\dagger(k) = \sqrt{\frac{1}{8\pi\hbar c k}} \left[ k A(k) - \frac{4\pi i c}{k} \Pi(k) \right]$$

NOTE:  $\beta, \beta^\dagger$  are linearly independent from  $\alpha, \alpha^\dagger$ .

together form a basis of the operator space since  $A(k), A^\dagger(k), \Pi(k), \Pi^\dagger(k)$  are needed to form field observables,

$$A_2(k) = A(k) + A^\dagger(k) \dots$$

Commutation rules:

$$[\alpha(k), \alpha(k')] = 0$$

$$[\alpha(k), \alpha^\dagger(k')] = \frac{1}{8\pi\hbar c k} \left[ -4\pi i c k (ik) + 4\pi i c k (-ik) \right] (2\pi)^3 \delta^3(k-k') = (2\pi)^3 \delta^3(k-k')$$

$$[\alpha(k), \beta(k')] = 0$$

$$[\alpha(k), \beta^\dagger(k')] = \frac{1}{8\pi\hbar c k} \left[ 4\pi i c k (ik) + 4\pi i c k (-ik) \right] = 0.$$

NOTE: we have used

$$[\Pi(n), A^\dagger(k)] = \left( [A(n), \Pi^\dagger(n)] \right)^\dagger = (i\hbar)^\dagger (2\pi)^3 \delta^3(n-n') = -i\hbar (2\pi)^3 \delta^3(n-n')$$

$\Rightarrow$  the set  $\{\alpha(n), \beta(n), \alpha^\dagger(n), \beta^\dagger(n)\}$  satisfies

Bosonic commutation rules as creation/annihilation operators of harmonic oscillators

NOTE:  $n$  is still in half-space.

$$\begin{aligned} \text{but naive extension } \alpha(-n) &= \sqrt{\frac{1}{\hbar \pi \omega}} \left[ \kappa A(-n) + \kappa \pi i c \Pi(-n) \right] = \\ &= \sqrt{\frac{1}{\hbar \pi \omega}} \left[ \kappa A^\dagger(n) + \kappa \pi i c \Pi^\dagger(n) \right] = \\ &= \beta(n) \end{aligned}$$

free-field Hamiltonian

$$H_f = \frac{1}{4\pi} \int \frac{d^3n}{(2\pi)^3} \left[ c^2 |\kappa \pi \Pi(n)|^2 + \omega^2 |A(n)|^2 \right]$$

can be rewritten as:

$$\begin{aligned} = \int \frac{d^3n}{(2\pi)^3} \frac{\hbar \omega}{2} \left[ \alpha^\dagger(n) \alpha(n) + \alpha(n) \alpha^\dagger(n) + \right. \\ \left. + \beta^\dagger(n) \beta(n) + \beta(n) \beta^\dagger(n) \right] \end{aligned}$$

where  $f$  is still in half-space!

Field operator

$$A(x) = \int \frac{d^3k}{(2\pi)^3} A(k) e^{ikx} = \int \frac{d^3k}{(2\pi)^3} (A(k) e^{ikx} + A(-k) e^{-ikx}) =$$

$$= \int \frac{d^3k}{(2\pi)^3} (A(k) e^{ikx} + A^\dagger(k) e^{-ikx}) = \dots$$

$$A(k) = \sqrt{\frac{2\pi\hbar c}{k}} (\alpha(k) + \beta^\dagger(k))$$

$$\dots = \int \frac{d^3k}{(2\pi)^3} \sqrt{\frac{2\pi\hbar c}{k}} \left[ (\alpha(k) + \beta^\dagger(k)) e^{ikx} + (\alpha^\dagger(k) + \beta(k)) e^{-ikx} \right] = \dots$$

Extending definition  $\alpha(-k) = \beta(k)$

$$\dots = \int \frac{d^3k}{(2\pi)^3} \sqrt{\frac{2\pi\hbar c}{k}} (\alpha(k) e^{ikx} + \alpha^\dagger(k) e^{-ikx}) = A(x)$$

With the same definition:

$$H_F = \int \frac{d^3k}{(2\pi)^3} \frac{\hbar c k}{2} [\alpha(k) \alpha^\dagger(k) + \alpha^\dagger(k) \alpha(k)]$$

NOTE: so far, we have always "forgotten" polarization degrees of freedom.

$$H_F = \int \frac{d^3k}{(2\pi)^3} \sum_{\epsilon \perp k} \frac{\hbar c k}{2} [\alpha(k, \epsilon) \alpha^\dagger(k, \epsilon) + \alpha^\dagger(k, \epsilon) \alpha(k, \epsilon)]$$

$$A(x) = \int \frac{d^3k}{(2\pi)^3} \sqrt{\frac{\hbar c}{2k}} \cdot \hat{\epsilon} \left[ \alpha(k, \epsilon) e^{ikx} + \alpha^\dagger(k, \epsilon) e^{-ikx} \right]$$

with  $\hat{\epsilon}$  unit vector  $\perp k$ .

Ground state

$|g\rangle$  is defined as  $\alpha(n)|g\rangle = 0$  for all  $n$ 's.

NOTE: proof of existence of  $|g\rangle$  can be found in  
C. Cohen-Tannoudji, B. Diu, F. Laloë, *Quantum Mechanics*,  
Vol. 1, chap. 5

All excited eigenstates characterized by  $n_n$ :

$$|e\rangle = \prod_n (\alpha(n))^{m_n} |g\rangle.$$

$\hookrightarrow$  mode  $k$  contains  $m_n$  photons!

Observables

$$B(\mathbf{r}) = \nabla \times A(\mathbf{r}) = \int \frac{d^3n}{(2\pi)^3} \sum_{\mathbf{e}, n} i \sqrt{\frac{2\pi\hbar c}{n}} (\mathbf{k} \times \hat{\mathbf{e}}) \cdot \left[ \alpha(n, \mathbf{e}) e^{i\mathbf{n}\cdot\mathbf{r}} - \alpha^\dagger(n, \mathbf{e}) e^{-i\mathbf{n}\cdot\mathbf{r}} \right]$$

$$\begin{aligned} \hat{E} &= -\frac{1}{c} \dot{A} = -\frac{1}{i\hbar c} [A, H] = \int \frac{d^3n}{(2\pi)^3} \sum_{\mathbf{e}} \sqrt{\frac{2\pi\hbar c}{n}} \frac{-1}{i\hbar c} \cdot \hbar c k \cdot \hat{\mathbf{e}} \cdot \left[ \alpha(n, \mathbf{e}) e^{i\mathbf{n}\cdot\mathbf{r}} - \alpha^\dagger(n, \mathbf{e}) e^{-i\mathbf{n}\cdot\mathbf{r}} \right] \\ &= \int \frac{d^3n}{(2\pi)^3} \sum_{\mathbf{e}} i \sqrt{\frac{2\pi\hbar c}{n}} n \cdot \hat{\mathbf{e}} \left[ \alpha(n, \mathbf{e}) e^{i\mathbf{n}\cdot\mathbf{r}} - \alpha^\dagger(n, \mathbf{e}) e^{-i\mathbf{n}\cdot\mathbf{r}} \right] \end{aligned}$$

NOTE: light-matter interaction term  $\frac{1}{2m} (P - \frac{q}{c} A)^2$  does not contribute to  $\hat{A}$ , nor to  $\hat{E}$ .

Commutation rules

$$\begin{aligned}
 [A_i(z), A_j(z')] &= \int \frac{d^3n}{(2\pi)^3} \int \frac{d^3n'}{(2\pi)^3} \sum_{\epsilon, \epsilon'} \frac{2\pi k c}{\sqrt{nn'}} \epsilon_i \epsilon_j \\
 &\cdot \left[ \alpha(n) e^{in z} + \alpha^\dagger(n) e^{-in z}, \alpha(n') e^{in' z'} + \alpha^\dagger(n') e^{-in' z'} \right] = \\
 &= \iint \sum_{\epsilon, \epsilon'} \frac{2\pi k c}{\sqrt{nn'}} \cdot \left[ \delta_{\epsilon\epsilon'} \delta(n-n') e^{in(n-z')} - \delta_{\epsilon\epsilon'} \delta(n-n') e^{-in(n-z')} \right] \cdot \epsilon_i \epsilon_j \\
 &= \int \frac{d^3n}{(2\pi)^3} \frac{2\pi k c}{-n} P_+(n) \underbrace{\left( e^{in(n-z')} - e^{-in(n-z')} \right)}_{\text{odd function in } n} = 0
 \end{aligned}$$

$$\begin{aligned}
 [E_i(z), A_j(z')] &= \int \frac{d^3n}{(2\pi)^3} \int \frac{d^3n'}{(2\pi)^3} \sum_{\epsilon, \epsilon'} \frac{i \sqrt{2\pi k c h}}{\sqrt{nn'}} \cdot \frac{\sqrt{2\pi k c}}{\sqrt{nn'}} \\
 &\cdot \epsilon_i \epsilon_j' \left[ \alpha(n, \epsilon) e^{in z} - \alpha^\dagger(n, \epsilon) e^{-in z}, \alpha(n', \epsilon') e^{in' z'} + \alpha^\dagger(n', \epsilon') e^{-in' z'} \right] \\
 &= i \frac{2\pi k c}{2\pi c} \int \frac{d^3n}{(2\pi)^3} \sum_{\epsilon} \epsilon_i \epsilon_j' \left[ e^{in(z-z')} + e^{-in(z-z')} \right] \\
 &= i \frac{2\pi k c}{2\pi c} \delta_{ij}^+ (z-z')
 \end{aligned}$$

$$[E_i(z), E_j(z')] = 0$$

Properties of the vacuum state :

zero-point (quantum) fluctuations

ground state  $|vac\rangle$  :  $\alpha(\mathbf{n}, \epsilon) |vac\rangle = 0$ .

$$\langle vac | E(\mathbf{r}) | vac \rangle = 0.$$

$$\text{but } \langle vac | E(\mathbf{r})^2 | vac \rangle = \int \frac{d^3n}{(2\pi)^3} \int \frac{d^3n'}{(2\pi)^3} \sum_{\epsilon \epsilon'}$$

$$(-) 2\pi\hbar c \sqrt{\hbar n'} \hat{E} \cdot \hat{E}' \langle vac | (\alpha(\mathbf{n}, \epsilon) - \cancel{\alpha(\mathbf{n}, \epsilon)}) \cdot (\cancel{\alpha(\mathbf{n}', \epsilon')} - \alpha(\mathbf{n}', \epsilon')) | vac \rangle$$

$$= \int \frac{d^3n}{(2\pi)^3} \sum_{\epsilon} 2\pi\hbar c k = \infty$$

reference of zero-point fluctuations diverges UV

Regularized form:  $\bar{E}(\mathbf{r}) = \int d^3r' g(\mathbf{r}-\mathbf{r}') E(\mathbf{r}')$

with  $g(\mathbf{r})$  localized around 0, but regular

$$\langle vac | \bar{E}(\mathbf{r})^2 | vac \rangle = \int d^3r d^3r' g(\mathbf{r}) g(\mathbf{r}') \cdot \int \frac{d^3n}{(2\pi)^3} \sum_{\epsilon} \int \frac{d^3n'}{(2\pi)^3} \sum_{\epsilon'}$$

$$2\pi\hbar c \cdot e^{i\mathbf{n}\cdot\mathbf{r}} e^{-i\mathbf{n}'\cdot\mathbf{r}'} \frac{1}{(2\pi)^3} \delta(\mathbf{n}-\mathbf{n}') \sum_{\epsilon \epsilon'} (\hat{E} \cdot \hat{E}')$$

$$= \int \frac{d^3n}{(2\pi)^3} 2\pi\hbar c \int d^3r g(\mathbf{r}) e^{i\mathbf{n}\cdot\mathbf{r}} \int d^3r' g(\mathbf{r}') e^{-i\mathbf{n}\cdot\mathbf{r}'}$$

$$= \int \frac{d^3n}{(2\pi)^3} 4\pi\hbar c |\tilde{g}(\mathbf{n})|^2$$

→ the  $\tilde{g}(k)$  function regularizes the integral and avoids UV divergence.

Similar strategy can be used to regularize light-matter Hamiltonian:

$$H_{\text{int}} = -\frac{q}{m} \vec{p} \cdot \vec{A}(0) = -\frac{q}{m} \vec{p} \cdot \int \frac{d^3k}{(2\pi)^3} \cdot \vec{E}(k, \epsilon) e^{i\mathbf{k}\cdot\mathbf{r}} + e^{i\mathbf{k}\cdot\mathbf{r}} + e^{i\mathbf{k}\cdot\mathbf{r}} e^{-i\mathbf{k}\cdot\mathbf{r}} \cdot \sqrt{\frac{2\pi\hbar c}{\omega}} \cdot \tilde{g}(k)$$

where we have defined  $\vec{A}(0) = \int d^3z g(z) \vec{A}(z)$

Quantum fluctuations of E, B, A fields are analogous to zero-point fluctuations of quantum harmonic oscillator:

$$X = \sqrt{\frac{\hbar}{2m\omega_0}} (a + e^{i\omega_0 t}), \quad P = \sqrt{\frac{m\hbar\omega_0}{2}} i(a^\dagger - a)$$

$$\langle g | X | g \rangle = 0, \quad \langle g | P | g \rangle = 0, \quad \langle g | P^2 | g \rangle = \frac{m\hbar\omega_0}{2} \langle g | i(a^\dagger - a) i(a^\dagger - a) | g \rangle$$

$$\langle g | X^2 | g \rangle = \frac{\hbar}{2m\omega_0} \langle g | (a + e^{i\omega_0 t}) (a + e^{i\omega_0 t}) | g \rangle = \frac{\hbar}{2m\omega_0} = \frac{m\hbar\omega_0}{2}$$

Zero-point energy  $\frac{\hbar\omega_0}{2}$  comes from:

$$\langle g | H | g \rangle = \langle g | \frac{p^2}{2m} + \frac{m\omega_0^2}{2} X^2 | g \rangle = \frac{\hbar\omega_0}{2}$$



$$\begin{aligned}
 H_F &= \int \frac{d^3n}{(2\pi)^3} \sum_{\epsilon} \frac{\hbar c n}{2} (a(n, \epsilon) a^\dagger(n, \epsilon) + a^\dagger(n, \epsilon) a(n, \epsilon)) \\
 &= \underbrace{\int \frac{d^3n}{(2\pi)^3} \sum_{\epsilon} \frac{\hbar c n}{2}}_{\text{zero-point energy } E_0} + \underbrace{\int \frac{d^3n}{(2\pi)^3} \sum_{\epsilon} \hbar c n a^\dagger(n, \epsilon) a(n, \epsilon)}_{\text{excitations above } |vac\rangle}
 \end{aligned}$$

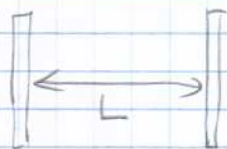
$E_0 \rightarrow E_0$  is UV-divergent.

$\rightarrow$  has no direct physical effect.

How to avoid infinite fluctuations in vacuum state:

1 - spontaneous emission is sometimes considered as a consequence of z.p.f.

2 - (static) Casimir effect:



- plane mirrors modify spectrum of z.p.f.

-  $E_0(L)$  is still UV-divergent, but the derivative

$$\frac{dE_0}{dL} \text{ is finite } \rightarrow \text{ finite force } F = - \frac{dE_0}{dL}$$

$\Rightarrow$  CASIMIR force

3 - Dynamical Casimir effect (DCE)

- fast motion of mirrors
- x.p.f. have no time to (adiabatically) follow the motion of mirrors and remain in ground state
- photons are emitted even in the absence of net currents

see e.g. :

- M. Kardar and R. Josthansen Rev. Mod. Phys. 71, 1233 ('99)
- A. Lenkecht, J. Opt. B: Quant. Sericlon. Opt. 7, 53 ('05)